

NALYSIS & RESULTS 2013-2020

NASA

C. R. LAWRENCE JET PROPULSION LABORATORY

BEYONDPLANCK CONFERENCE 2020 NOVEMBER 18

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- Goal: measure the temperature anisotropies of the cosmic microwave background (CMB) to fundamental limits down to 5', also measure polarization better than ever before
 - Two state-of-the-art cryogenic instruments
 - Nine bands, 30 GHz to 857 GHz. 30–353 GHz polarized.





PAPER AND DATA RELEASES

	Data [Months]									
						Stokes		CALIBRATION		
PAPERS	Release	LFI	HFI	TOD	Maps	LFI	HFI	Solar	Orbital	Science
Early		10	10	no	yes	Ι	Ι	yes		astrophysics
2013	PR1	15.5	15.5	yes	yes	Ι	Ι	yes		astro+cosmo
2015	PR2	50.5	29	yes	yes	IQU	$I(QU)^{\mathrm{a}}$		yes	cosmology
2018	PR3	50.5	29	yes	yes	IQU	$(I)^{\mathrm{b}}QU$		yes	cosmology
NPIPE	PR4	50.5	29	yes	yes	IQU	IQU	•••	yes	limited

^a Not usable at low ℓ .

^b Not changed from 2015.



PLANCK COLLABORATION PAPERS

		-	# of Pages			
	# of					
\mathbf{Set}	PAPERS	Total	per Paper	Mean		
Early results	26	584	7-56	22.5		
2013 results	32	992	20-66	31.0		
2015 results	28	1067	17 - 99	38.1		
2018 results	12	661	31 - 95	57.3		
Intermediate	57	1311	12-88	23.0		
Planck/Bicep	1	17				
Total	156	4632		29.7		

- 48,779 citations as of 17 November 2020
- 8,962 citations for Parameters 2015
- Another >226 papers by Planck team members that are not Planck Collaboration papers





- For the 2013 results papers, the CMB power spectra produced by LFI and HFI agreed over the first peak extremely well
 - But disagreed with WMAP by about 2%, with concerning significance
 - This was a puzzle for a while
- *Planck* 2013 results. XXXI. Consistency of the *Planck* data started to sort it out
 - For LFI, the full 4π beam response used in the dipole calibration needed improvement (~ 0.2% in power spectra)
 - For HFI, near sidelobe estimates (from "main" beam out to 5°) needed improvement (beam window function corrections of 0.5–1.2%). Time constants were a factor.
 - Planck in 2013 was calibrated on the WMAP solar dipole, not the orbital dipole. WMAP, on the other hand, was calibrated on the orbital dipole WMAP solar dipole errors fed directly into Planck errors.
 Foreground errors in the WMAP solar dipole fed directly into Planck errors, but didn't affect WMAP.
 - Corrections for foregrounds, including unresolved sources, were necessary
 - LFI and HFI agreed well within uncertainties. Planck WMAP was a 1.5–2 σ difference
- Overall 2013 Planck calibration errors were estimated at 0.62% for 70 GHz maps and 0.54% for 100 and 143 GHz maps
- Planck 2018 and NPIPE are better than 0.1%. Stay tuned for later talks.
 - The improvement was a result of many, many factors.





- This is the hard part, of course!
- Planck ultimately was not limited by systematics except at the lowest few multipoles
- The hardest to deal with was an HFI combination of
 - ADC non-linearities (not anticipated)
 - Way out-of-spec EMC from the 4-K-cooler drive electronics (known before launch)
 - Limited downlink bandwidth
 - 40 successive readout samples were summed on board, so sample-by-sample correction was impossible
- Second-hardest was another HFI combination, of
 - Beams
 - Time constants
 - 'Long' cosmic rays
- There are no shortcuts in finding and mitigating systematic problems
 - Must connect any effect to its hardware cause
 - Must be able to reproduce it with simulations built on the real behavior of hardware
 - Sometimes valuable techniques such as blind analysis and machine learning don't help



THE UNIVERSE: TEMPERATURE, NINE FREQUENCIES



PLANCK POLARIZATION, SEVEN FREQUENCIES







SEPARATION OF 'FOREGROUNDS' FROM THE CMB



Planck 2018 results. I.

Temperature

- All components smoothed to 1°
- Sky fractions 81–93% of sky

Polarization

- All components smoothed to 40'
- Sky fractions 27–93% of sky

CMB AND FOREGROUND STOKES I MAPS





CMB AND FOREGROUND STOKES Q, U MAPS













(A few % of the sky in the plane of the Milky Way is filled in with a "constrained realization".) (Regions inside the grey lines aren't used for cosmology, as residual foreground emission is too high.)



A "SIMPLE" 6-PARAMETER ΛCDM MODEL FITS THE PLANCK DATA EXTREMELY WELL!

The TT, TE, EE, and CMB lensing spectra Are consistent with each other under the Assumption of the base Λ CDM cosmology.



Assumptions Underlying ΛCDM

- 1 Physics is the same throughout the Universe
- 2 General Relativity is an adequate description of gravity
- 3 On large scales, the Universe is statistically the same everywhere
- 4 The Universe was once hot and dense, and has been expanding
- 5 Five basic cosmological constituents
 - Dark energy that behaves like the energy density of the vacuum
 - Dark matter that is pressureless, stable, and non-electromagnetically-interacting
 - Regular "atomic" (baryonic) matter that behaves just like it does on Earth
 - The photons we observe as the CMB
 - Neutrinos that are almost massless, and stream like non-interacting, relativistic particles
- 6 The curvature of space is very small
- 7 Variations in density were created early, and are Gaussian, adiabatic, and nearly scale invariant, as predicted by inflation
 - I.e., Proportional in all constituents, and with similar amplitudes as a function of scale
- 8 The Universe has "trivial" topology (in particular, it is not periodic or multiply connected)







- $A_{\rm s}$, $n_{\rm s}$ inflation fluctuations; 10^{-35} s;
- $\Omega_{\rm b}h^2$, $\Omega_{\rm c}h^2$ baryons and cold dark matter; first few minutes
 - 0.6% and 0.8% precision
- $\theta_{\rm MC}$ sound horizon; 370,000 years
 - 0.03% precision
- τ reionization optical depth; 13.8 billion years



ANGULAR POWER SPECTRUM + BEST-FIT MODEL





POLARIZATION SPECTRA, SAME MODEL



Red line is again the "best-fit" model to the temperature data.

• Spectacular agreement between inferences from temperature and polarization

LENSING POTENTIAL — ALL THE MASS IN THE UNIVERSE



• Lensing now measured at 40σ .



ΛCDM Model Parameters



Parameter	Planck + BAO	N_{σ}
$\Omega_{ m b} h^2 [18.79 { m yg} { m m}^{-3}]$	0.02242 ± 0.00014	160
$\Omega_{ m c} h^2 [18.79{ m ygm^{-3}}]$	0.11933 ± 0.00091	131
$100\theta_{\rm MC}$	1.04101 ± 0.00029	3590
au	0.0561 ± 0.0071	7.9
au (NPIPE)	0.051 ± 0.006	8.5
$\ln(10^{10}A_{\rm s})$	3.047 ± 0.014	218
<i>n</i> _s	0.9665 ± 0.0038	254

$H_0[\rm kms^{-1}Mpc^{-1}]$	67.66 ± 0.42	161
Ω_{m}	0.3111 ± 0.0056	55
$z_{\rm reionization}$	7.82 ± 0.71	11
$z_{ m recombination}$	1059.94 ± 0.30	3533
Age[Gyr]	13.787 ± 0.020	689

68% confidence limits





CONSISTENCY WITH OTHER DATA

- Baryon Acoustic Oscillations (BAO; distance scale)
- Primordial nucleosynthesis
- Type la supernovae
- Direct measures of H_0
- Redshift-space distortions
- Rich clusters of galaxies





Vertical axis is the ratio of the acoustic-scale distance as determined by the experiments plotted to that determined by Planck. $D_V(z)/r_{drag}$ is the acoustic-scale distance ratio

 $r_{\rm drag} = comoving$ sound horizon at end of baryon drag epoch

$$D_V = \left[(1+z)^2 D_A^2(z) \frac{cz}{H(z)} \right]^{1/3}$$

$$D_A = \text{angular diameter distance}$$

BIG BANG NUCLEOSYNTHESIS



The width of the green stripes corresponds to 68% uncertainties in nuclear reaction rates and on the neutron lifetime. The horizontal bands show observational bounds on primordial element abundances compiled by various authors, and the red vertical band shows the Planck TT+lowP+BAO bound on $\Omega_{\rm b}h^2$ (all with 68% errors). The BBN predictions and CMB results shown here assume $N_{\rm eff} = 3.046$ and no significant lepton asymmetry.



The standard Λ CDM model fits really well. Do models with more parameters fit better?

- Ω_{K} (curvature)
- Σm_{ν} (neutrino mass) and $N_{\rm eff}$ (effective number of "neutrino" species)
- $Y_{\rm P}$ (helium fraction)
- $dn_{\rm s}/d\ln k$ (curvature in the input fluctuation spectrum)
- $r_{0.002}$ (tensor-to-scalar ratio at $k_0 = 0.002 \,\mathrm{Mpc^{-1}}$)
- w (dark energy equation of state, constant)
- Huge number of variations posted on the archive.

NO COMPELLING EVIDENCE FOR ANY OF THESE EXTENSIONS



 CMB + later-time data from lensing and BAO lead to remarkable constraints on spatial curvature...

 $\Omega_k = 0.0007 \pm 0.0019(95\%)$

Non-Gaussianity



- $f_{\rm NL}^{\rm local} = -0.9 \pm 5.1$
- $f_{\rm NL}^{\rm equil} = -26 \pm 47$
- $f_{\rm NL}^{\rm ortho} = -38 \pm 24$

INFLATION CONSTRAINTS FROM PLANCK + BAO



- $V(\phi) \propto m^2 \phi^2$ and $\lambda \phi^4$ excluded. Models with $r \sim (1 n_s)^2$ (concave potentials) are consistent with data.
- \Rightarrow Energy scale of inflation $V_* = (1.94 \times 10^{16} \,\text{GeV})^4 r_{0.002}/0.12)$





 From Planck 2018 VI, combining Planck data with BICEP-Keck 2015 data (BICEP2 Collaboration 2018), we get

 $r_{0.002} < 0.06, \quad 95\%$

 From a new paper under review at A&A, combining NPIPE-processed Planck data with BICEP-Keck 2015 data, we get

 $r_{0.002} < 0.044, \quad 95\%$

Astronomy & Astrophysics manuscript no. planck_tensor November 17, 2020

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Planck constraints on the tensor-to-scalar ratio

M. Tristram¹, A. J. Banday^{2,3}, K. M. Górski^{4,5}, R. Keskitalo^{6,7}, C. R. Lawrence⁴, K. J. Andersen⁸, R. B. Barreiro⁹, J. Borrill^{6,7}, H. K. Eriksen⁸, R. Fernandez-Cobos⁹, T. S. Kisner^{6,7}, E. Martínez-González⁹, B. Partridge¹⁰, D. Scott¹¹, T. L. Svalheim⁸, H. Thommesen⁸, and I. K. Wehus⁸

- ¹ Université Paris-Saclay, CNRS/IN2P3, IJCLab, 91405 Orsay, France
- ² Université de Toulouse, UPS-OMP, IRAP, F-31028 Toulouse cedex 4, France
- ³ CNRS, IRAP, 9 Av. colonel Roche, BP 44346, F-31028 Toulouse cedex 4, France
- ⁴ Jet Propulsion Laboratory, California Institute of Technology, 4800 Oak Grove Drive, Pasadena, California, U.S.A.
- ⁵ Warsaw University Observatory, Aleje Ujazdowskie 4, 00-478 Warszawa, Poland
- ⁶ Computational Cosmology Center, Lawrence Berkeley National Laboratory, Berkeley, California, U.S.A.
- ⁷ Space Sciences Laboratory, University of California, Berkeley, California, U.S.A.
- ⁸ Institute of Theoretical Astrophysics, University of Oslo, Blindern, Oslo, Norway
- ⁹ Instituto de Física de Cantabria (CSIC-Universidad de Cantabria), Avda. de los Castros s/n, Santander, Spain
- ¹⁰ Haverford College Astronomy Department, 370 Lancaster Avenue, Haverford, Pennsylvania, U.S.A.
- ¹¹ Department of Physics & Astronomy, University of British Columbia, 6224 Agricultural Road, Vancouver, British Columbia, Canada

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- 6-parameter Λ CDM fits!
- CMB and large-scale structure (BAO) provide a consistent picture
- Extensions of the model have been tested. They are not required, and provide many limits
- Primordial fluctuations are to good approximation
 - Isotropic (but not completely)
 - Gaussian
 - Nearly scale invariant ($n_{\rm S} = 1$ is excluded at 9σ)
- Space is flat
- Matter is mostly dark and cold
- Baryon density, helium and deuterium abundances, are consistent with BBN
- There are no gravitational waves at the < 10% level





- We don't know what dark matter is.
- We don't have any idea at all what dark energy (Λ) is.
- The physics of inflation are largely unknown



- Planck met its goals
- The Planck data releases and papers are a major and wonderful contribution to our knowledge of the Universe
- There is much left to do with the Planck data...
- ...and a lot left to squeeze out of the CMB



